XENON100

PAUL SCOVELL ON BEHALF OF THE XENON100 COLLABORATION DM2012 - 24/02/12



XENON100 Collaboration





XENON100



• Compared with XE10:

• x10 fiducial mass increase

• x100 background decrease

• All components screened at dedicated screening facilities at LNGS

•161kg LXe total, 62kg target volume, 99kg LXe veto. 15cm Radius, 30cm drift

• Cathode at -16 kV, drift field of 0.533 kV/cm. Proportional scintillation region with field ~12 kV/cm.

• Custom-made low radioactivity HV feedthroughs.

Aprile *et al.*, Astropart.Phys.35:43-49,2011



XENON100 WAVEFORMS





FIRST RESULTS FROM XE100



After unblinding and application of a new noise removal cut, 3 events remain with expectation of 1.8±0.6.
Uniform in r and z.

Consistent with background leakage

100.9 DAY RESULTS FROM XE100



7x10⁻⁴⁵ cm² at 50 GeV/c² Aprile *et al.*, Phys. Rev. Lett. **107**, 131302, 2011



UPDATE SINCE 100.9 DAY RUN

After analysis of 100 days of data:

- No changes made to TPC detector not opened between runs
 - Improved trigger
 - Able to probe to lower energies
- Reduction of Krypton background
 - Circulation of Xe through dedicated distillation column
- Improved methods to tackle anomalous background
 - Multiple scintillation, single ionisation
 - Electronic noise
- Improved statistics in electron and nuclear recoil calibrations
 - Now using regular Co-60 and Th-232 measurements
 - Cs-137 still taken regularly to monitor electron lifetime. Lifetime now up to ~0.5ms



TRIGGER IMPROVEMENTS

• 100% efficiency down to 150 PE in S2



• Improvement in trigger electronics and increased electron lifetime allows extension of reach to lower energies



BACKGROUND REDUCTION

Reduction of intrinsic background - Kr85

- Kr 85 is internal background
- Decays via electronic process, indistinguishable from gamma background
- 100 days run suffered in sensitivity due to increased Kr/Xe from accidental leak
- In Autumn 2010, xenon circulated through dedicated distillation column
- Kr/Xe reduced significantly with respect to 100.9 days run
- Dedicated measurement gives (19±1)
 PPT





BACKGROUND REDUCTION

Removal of anomalous background

Pattern likelihood algorithm applied compares S1 light pattern to expected Trained using Monte-Carlo simulation





CALIBRATIONS - GAMMAS

Major component of background comes from gamma-ray emission. Sources used to model expected background.





CALIBRATIONS - NEUTRONS





CALIBRATIONS - NEUTRONS





L_{eff} input 2.26±0.03 PE/keVee



LEFF

• Energy scale (E_{nr}) is set using the S1 signal

 $E_{\rm nr} = \frac{S1}{L_{y,\rm er}} \frac{1}{\mathcal{L}_{\rm eff}(E_{\rm nr})} \frac{S_{\rm er}}{S_{\rm nr}}$

Light yield (L_{y,er}) is the light yield for electron recoils of 122keVee

• Snr and Ser represent quenching factors due to application of drift field

• Relative scintillation efficiency given as:



 $\mathcal{L}_{\text{eff}}(E_{\text{nr}}) = \frac{L_{y,\text{er}}(E_{\text{nr}})}{L_{y,\text{er}}(E_{\text{ee}} = 122 \text{ keV})}$

Plante *et al.*, Phys. Rev. C **84**, 045805, 2011



DETECTOR STABILITY



Detector has been filled since 2010. Temperature plot shows fantastic stability outside of chiller maintenance



DATA STATUS



Dark matter run started March 2011, ~210 live days accumulated
Much more ER calibration data, ~25 live days 60Co and ~25 live days 232Th



DATA STATUS THUS FAR



No noise events, all below 2PE are good

66% background reduction with veto cut

30 kg Fiducial - 198 days

30 kg Fiducial - 100.7 days



Significant reduction in background rate between the two runs

Background now comparable with that of 11.17 day result E.Aprile *et al.*, Phys. Rev. Lett. **105**, 131302, 2010



PROJECTED REACH



Limit assumes 200 day run background free and all acceptances included

Blind analysis underway Result in the spring



CONCLUSIONS

- Detector cold and filled > 1 year
 - 210 live days of dark matter
- Kr background reduced to low level
- Improvements in calibration techniques
- Result due in the spring
- Operation of XENON100 will continue through the XENON1T installation at LNGS