

# The EDELWEISS DM search Phase II to Phase III

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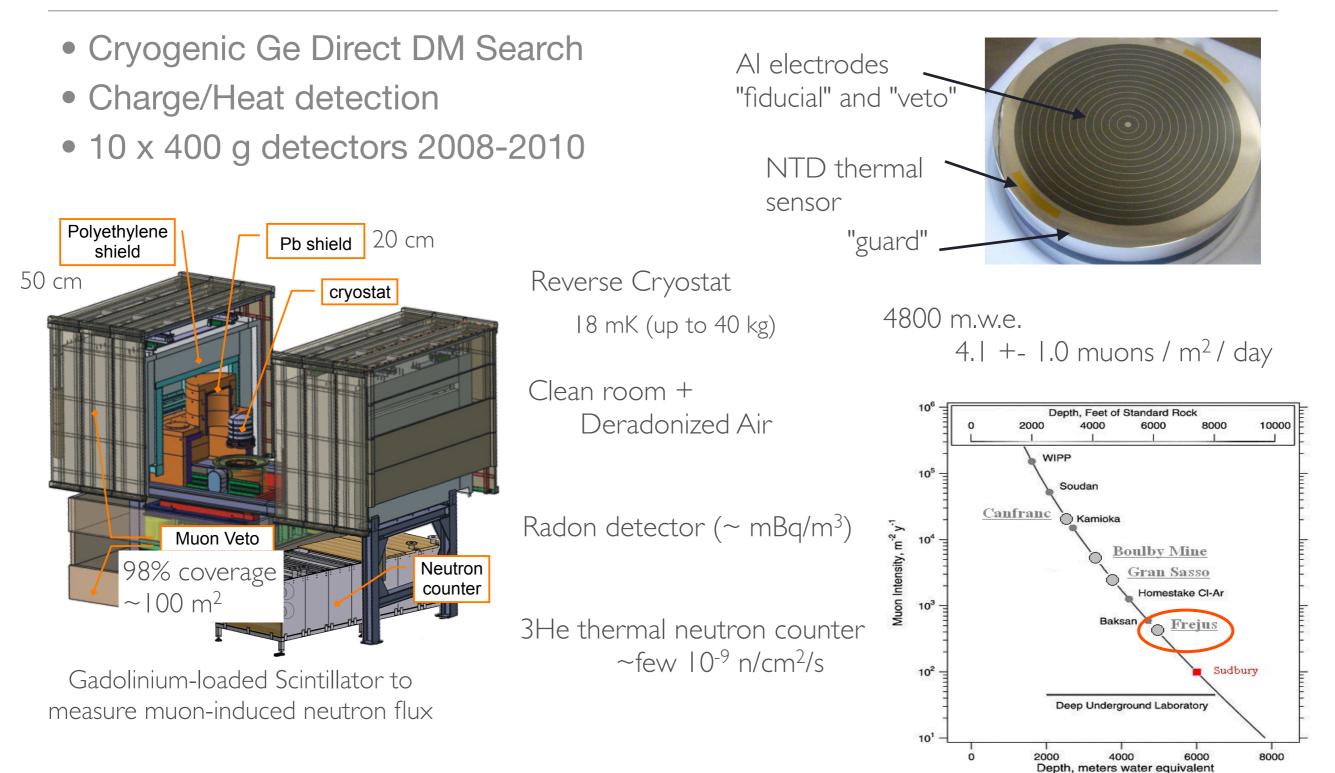
on behalf of the EDELWEISS Collaboration

CEA, Saclay (IRFU and IRAMIS) IPNL (CNRS/IN2P3 and Université de Lyon) CSNSM Orsay (CNRS/IN2P3 and Université Paris-Sud) Neél Grenoble (CNRS/INP) KIT (IK / EKP/ IPE), Karlsruhe JINR, Dubna University of Oxford University of Sheffield LSModane (CEA/CNRS)



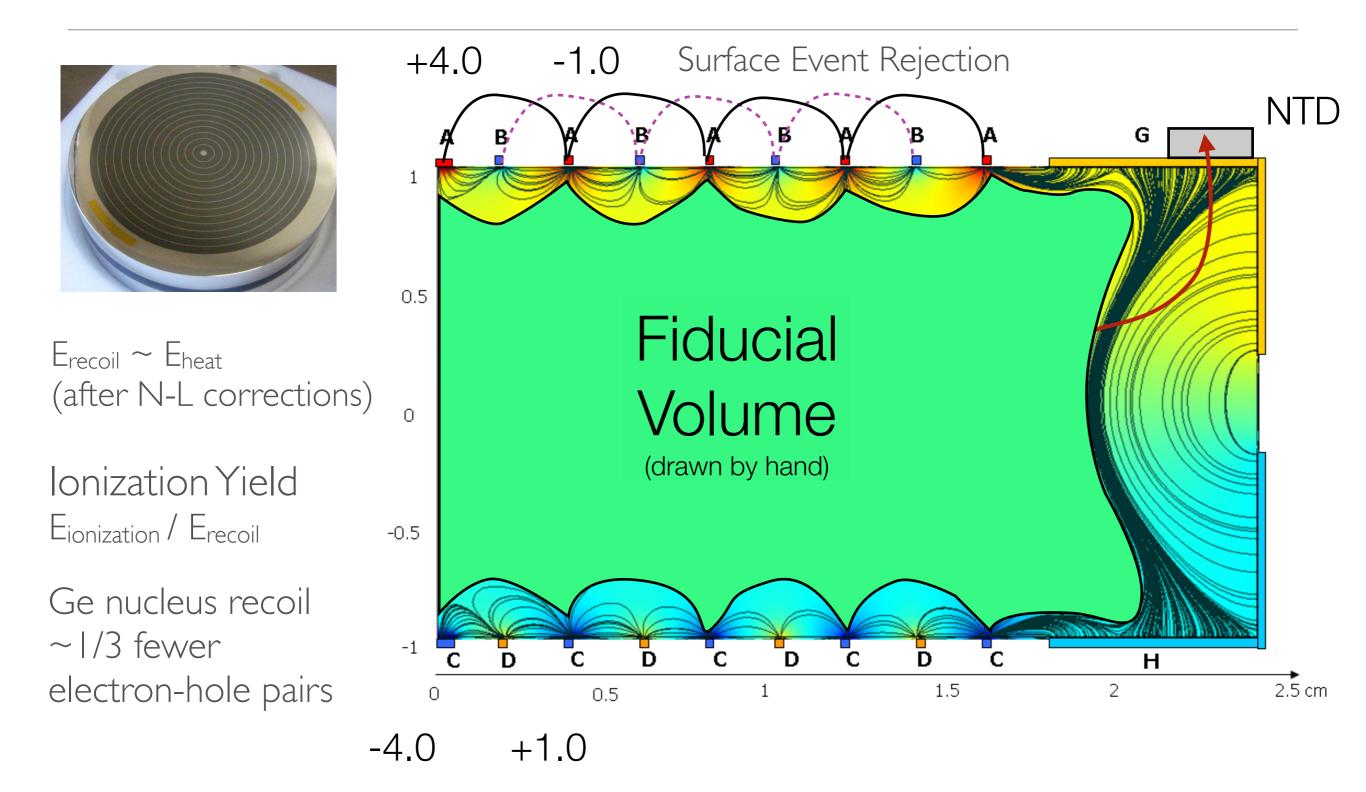


# Edelweiss II at LSM (Frejus Tunnel)





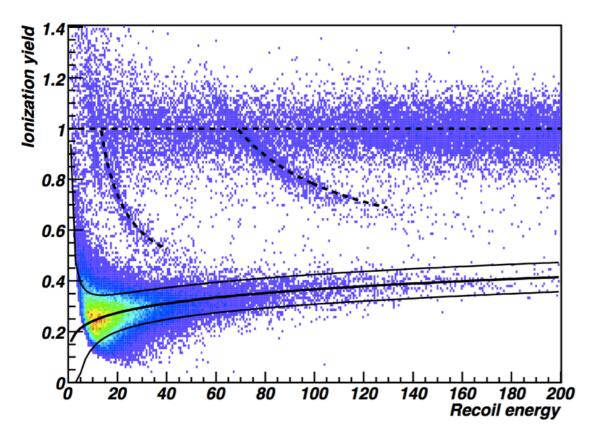
### An Edelweiss Detector





## Gamma Band and Nuclear Recoil Band

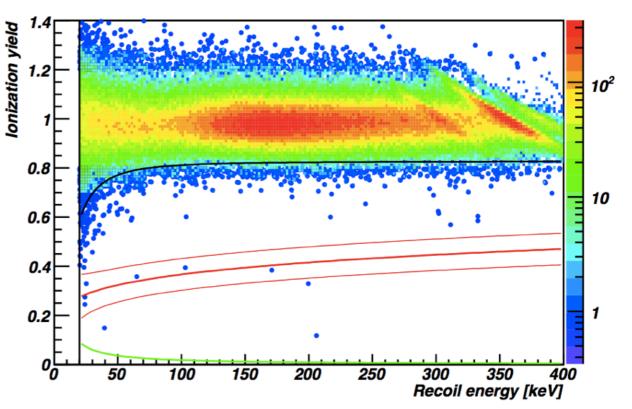
neutrons



Ge recoil lon yield =  $0.16 E_{rec}^{0.18}$ 

P. Di Stefano, et al., Astropart. Phys. 14 (2001) 329.O. Martineau, et al., Nucl. Instrum. Meth. A 530 (2004) 426.

<sup>133</sup>Ba (347k events)



1 NR for every 30k gammas between 20 and 200 keV

A. Broniatowski, et al., Phys. Lett. B 681 (2009) 305.

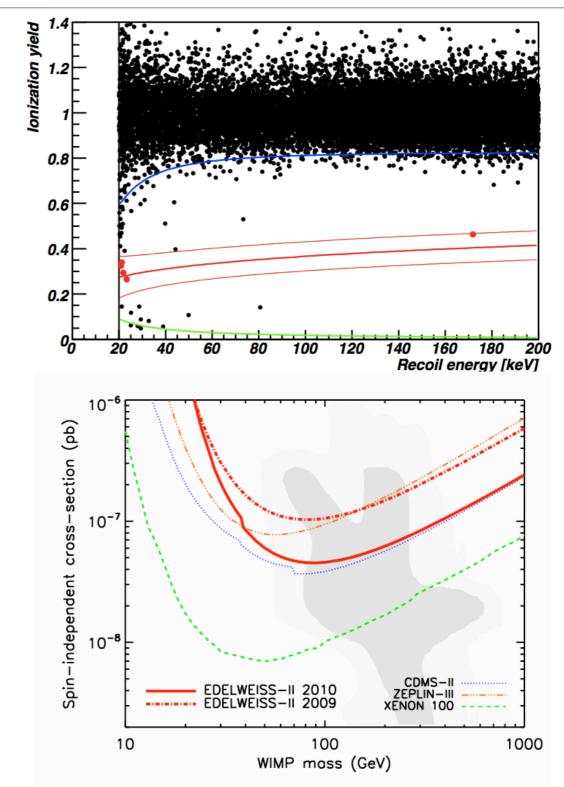


# Edelweiss II Results. Data from 2008 - 2010

- Ionization Yield (Eion / Erecoil) v Erecoil
- July Nov 2008 and April 2009 May 2010
- 427 kg \* days
- 384 kg \* days in 90% NR band
- 5 events in NR band

WIMP Halo local density of 0.3 GeV/c<sup>2</sup> Maxwellian velocity distribution  $v_{rms} = 270$  km/s  $v_{Earth} = 235$  km/s  $v_{escape} = 544$  km/s

$$\sigma_{SI} < 4.4 \times 10^{-8} \text{ pb at } 90\% \text{ CL for}$$
  
 $M_{WIMP} = 85 \text{ GeV/c}^2$ 

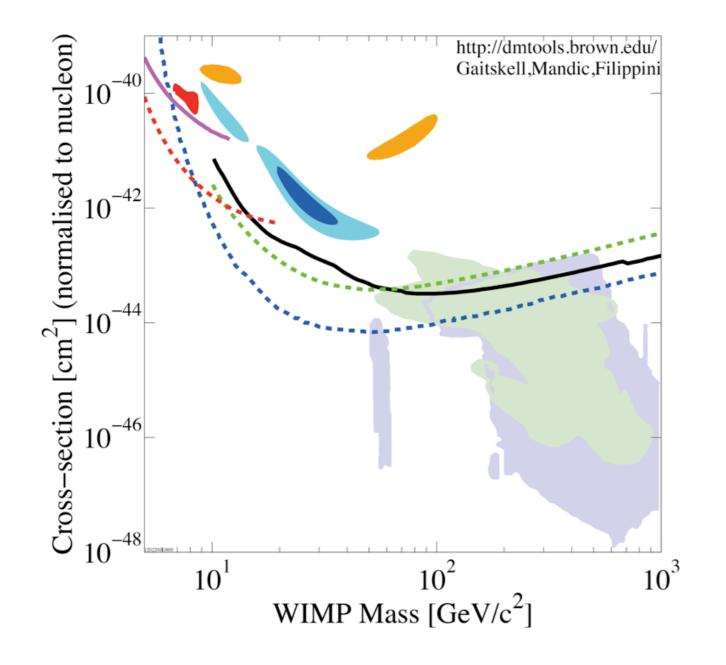


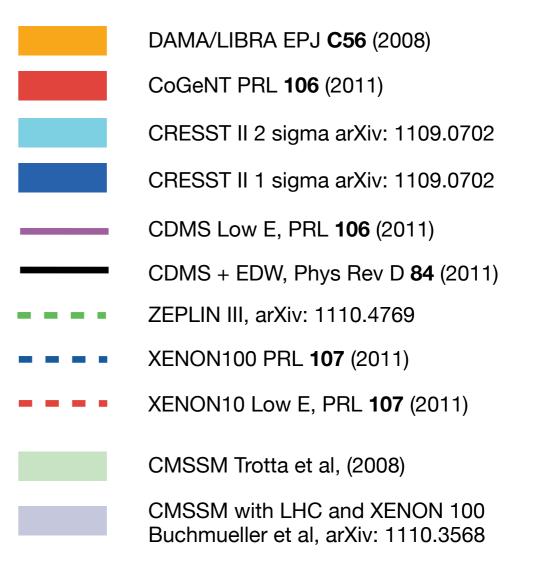
E. Armengaud, et al., Phys. Lett. B 702 (2011) 335–329.

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# **Direct Detection Current Status**



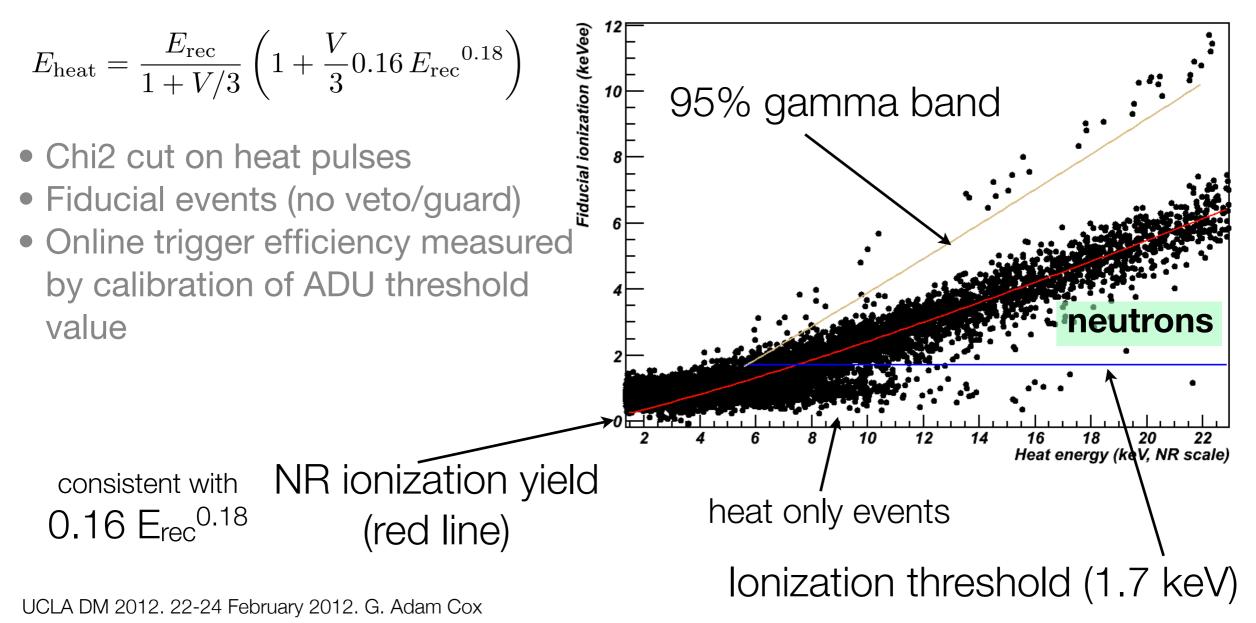




# **Preliminary** Low Energy Analysis of ID3

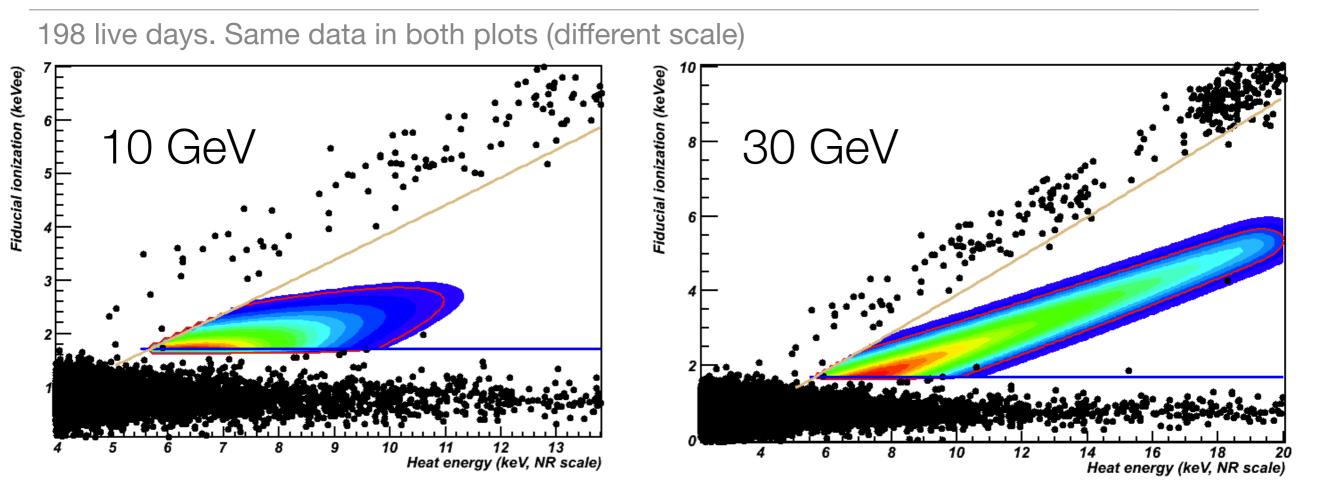
- Search below 20 keV
- Neutrons to 5 keV
- Energy based on heat only

Strong (10<sup>5</sup> n/s) AmBe neutron calibration 2010 placed outside of all shielding and muon veto





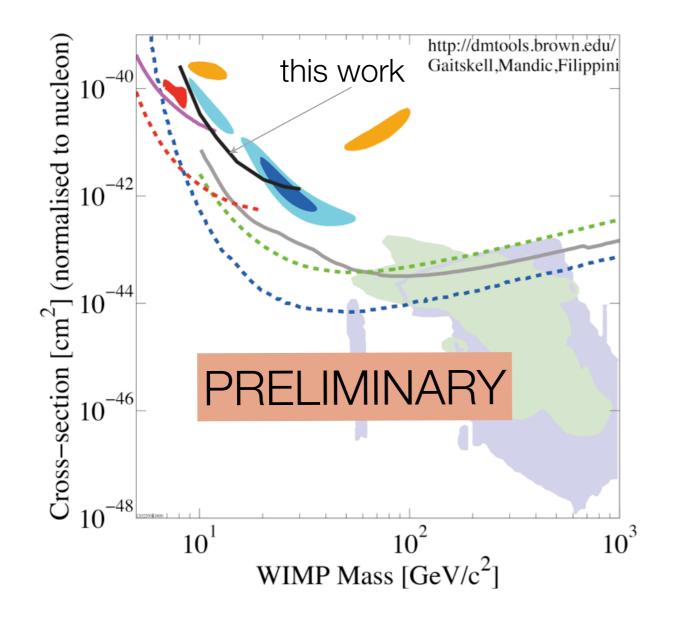
### WIMP search : ID3 low-energy background data



- Compute the density of a WIMP signal for a given mass (using measured resolutions and efficiencies) : colour contours
- Count the WIMP candidates within a ROI fixed to have 90% efficiency to the WIMP signal (red contour)
- No convincing signal (from 1 to 3 candidates depending on  $M_X$ ); derive 90% CL limit on  $\sigma_{SI}$  from Poisson statistics
- Estimated backgrounds:
  - ~ 0.25 events from heat-only pulses (using a background model)
  - $\sim 0.5$  events from fiducial gamma-rays
  - ~0.5 neutrons
  - No significant surface event background (compare to CDMS) !!!



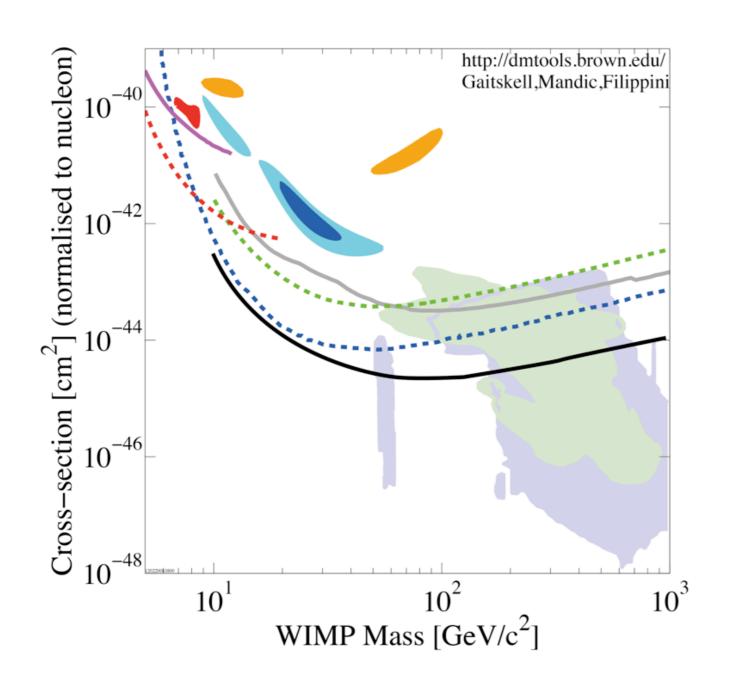
# Preliminary low-mass WIMP limit ~ 30 kg.d



- Combination with detectors of equivalent performance (in progress)
- Limited by Heat Only events
- ID3 detector (this analysis) has best resolution
- Poisson limits (optimal interval to come)
- (2 NTDs per bolometer in EDW III)



# EDELWEISS III



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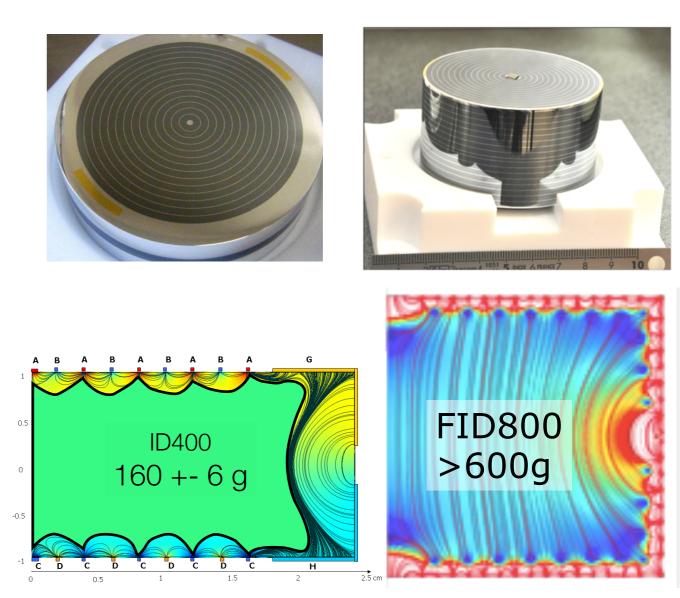
### Goals

- 3000 kg day exposure
- Explore low mass region
  - consistently < 1 keV FWHM ionization resolution
- Reduce background by factor of 10
  - shielding, material selection
  - further quantify gamma/surface rejection



### Increased Exposure: More detectors and Fiducial Volume

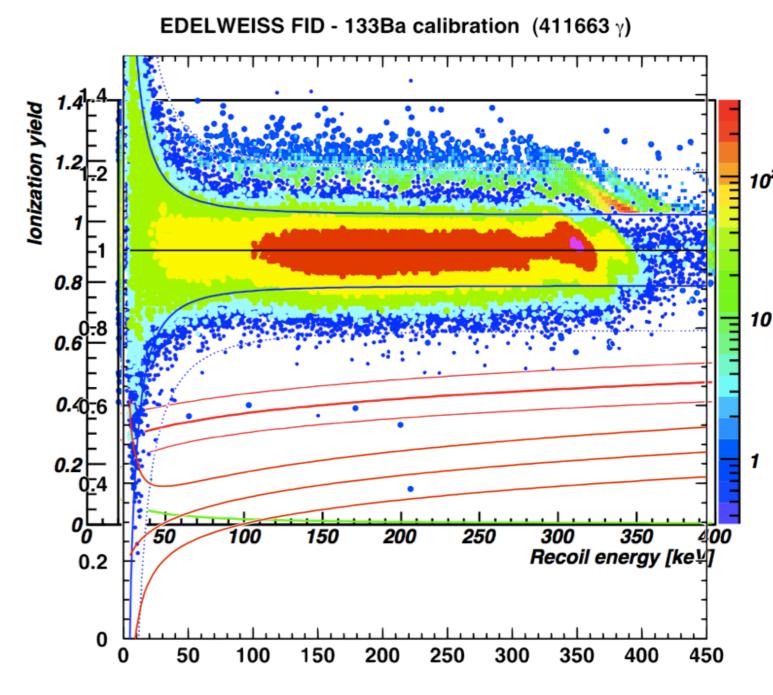
- Inter-Digitized -- Fully Inter-Digitized
- ~40% -- > 75%
- 400g -- 800g
- 40 new detectors
- 150 live days = 3000 kg days
- ~1 detector fabricated per week
- ~20 detectors ready by summer
- Expected full delivery Fall 2012





## Gammas

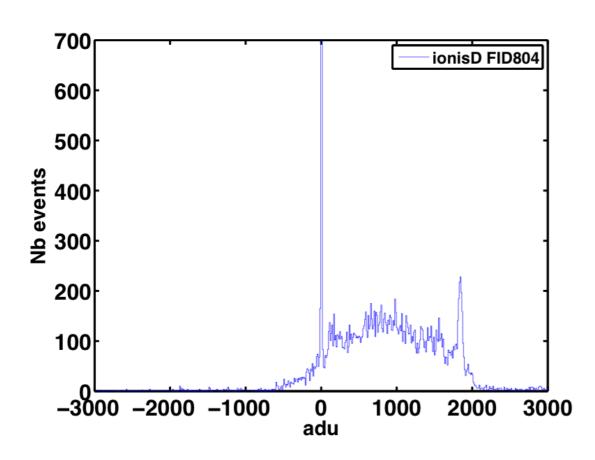
- Run12 ID: 6 events in NR band out of 300k events
- FIDs show better gamma rejection (400k events)
- Low Field regions on edges of ID detectors eliminated by FID

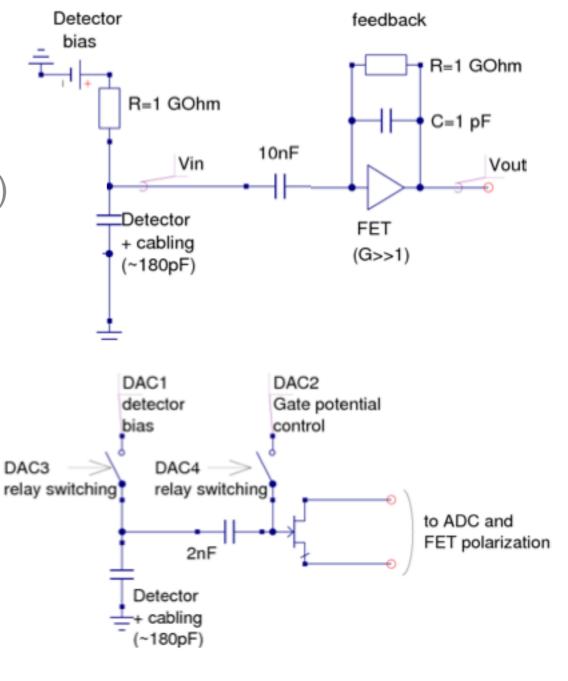




# Efforts to improve Energy Resolution

- New Front-End Electronics (in progress)
- No Feedback Resister on FET
- Relays recharge electrodes, reset FET
- < 1 keV FWHM baseline resolution (500 eV)







# Backgrounds in Edelweiss II

#### Ambient Neutrons:

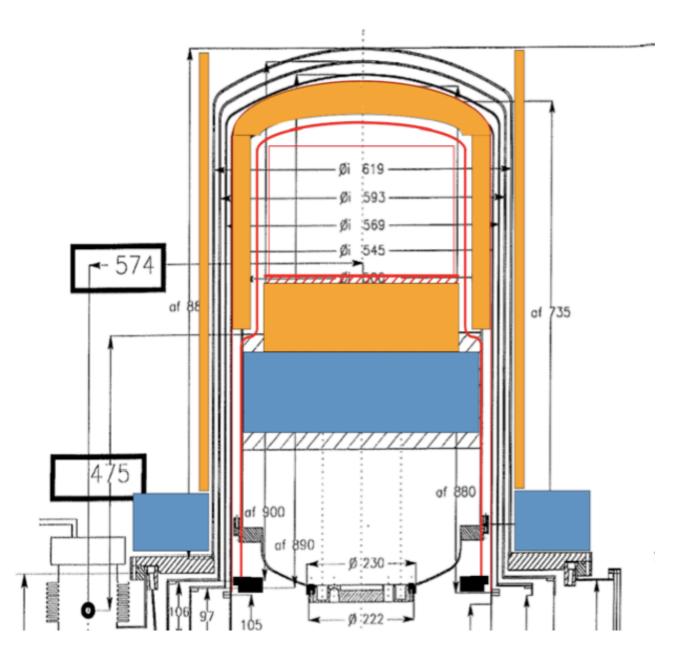
	Material	Mass (kg) or	U/Th	# Events	
		thickness (cm)	concentrations		
Hall walls	Concrete	30 cm	1.9 / 1.4 ppm	< 0.10	Radiopurity measurements +
Hall walls	Rock	$300 \mathrm{~cm}$	$0.8 \ / \ 2.5 \ \mathrm{ppm}$	< 0.01	GEANT4 Simulation
Shielding	Lead	20  cm (40  tonnes)	< 0.01 ppb U	< 0.08	of neutron background
Shielding	Polyethylene	50 cm	0.4 / < 0.5  ppb	< 0.05	of field off background
Support	Stainless steel	$0.6 \text{ cm} (\sim 200 \text{ kg})$	0.4 / < 1  ppb	< 0.01	
Support	Mild steel	8.6 tonnes	0.4 / < 1  ppb	< 0.04	- Measured neutron spectrun
Cryostat	Copper	$\sim 500 \text{ kg}$	< 0.3 / < 0.5  ppb	< 0.03	
Electrodes	Aluminium	< 0.03 g	< 1000 ppb U/Th	< 0.01	- SOURCES4A spectrum
Crystal holders	PTFE	$\sim 20 \text{ g}$	< 0.2 / < 0.5  ppb	< 0.01	
Connectors Cables	Al, plastics PTFE	$0.32~{ m kg} \ pprox 1.35~{ m kg}$	170 / 110 ppb 0.8 / < 1.5 ppb	< 0.40 < 0.66	
Muon-Ind	uced Neutron	s (not rejected by	Muon Veto)	< 0.4	measured
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	uced Neutron utron Backgro		Muon Veto)	< 0.4	measured
Total Ne	utron Backgro		Muon Veto)	<  .8 < 0.9	measured
Total Nei Gamma (	utron Backgro Contaminatic	ound		<  .8	

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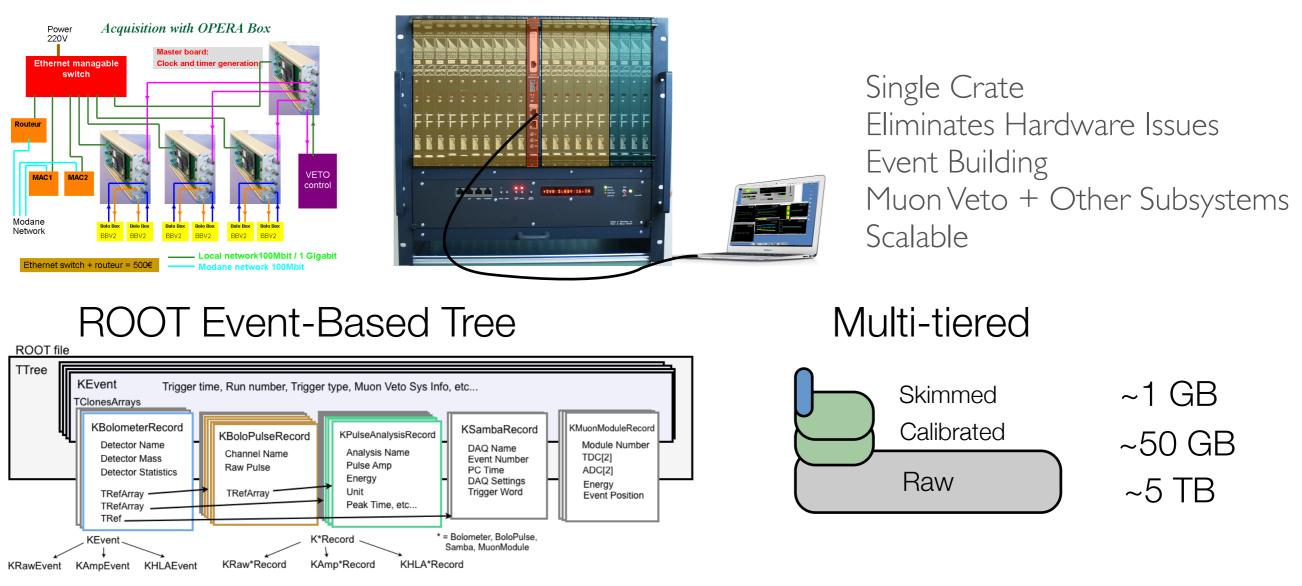
# Neutron Shielding

- Factor 6 reduction in neutrons
- Redesigned Cryostat, internal detector galette.
- Block neutrons from Pb Shield and cables/connectors
- Add Inner PE Shielding
- PE Radiopurity Measurement
- Material selection improvements for cables and connectors
- Cryostat design to be finalized in next month.





# Data Management for 40 x 6 channels



Database Driven DAQ and Data Process Monitoring System

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DAQ Conditions Metadata Data Processing Records Data location





## Conclusions

- EDW II Run12 low energy analysis to proceed detector by detector. Initial results indicate sensitivities to masses lower than previously reported by EDW.
- EDW III ....
  - Fabricating a new detector each week: Fall 2012 full 40-detector array
  - Redesigning cryostat / cables / shielding / electronics
  - Optimizing pulse amplitude estimate algorithms
  - Data management in place to handle expected large data set.
  - Currently taking gamma calibration data to further quantify gamma rejection capabilities
  - Characterizing NTDs. Two NTDs per detector to study heat only events.
  - Plan to take data on ~20 detectors starting early this summer

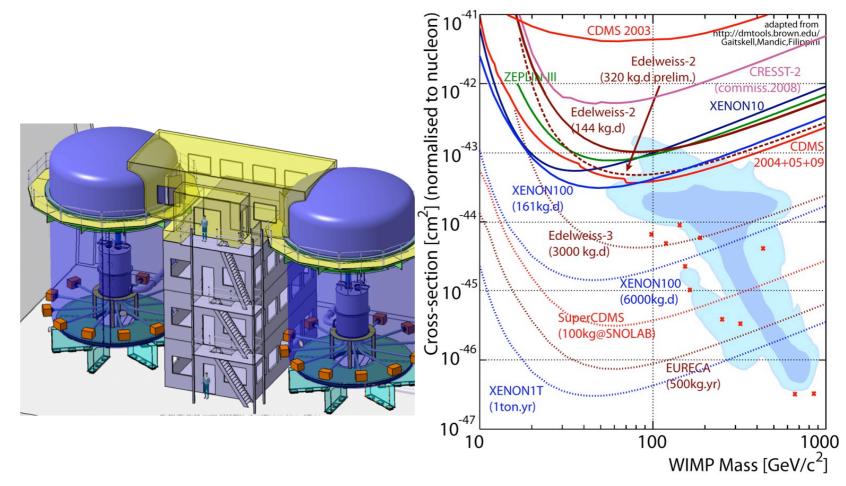


# Outlook: EURECA

Multi-target (Ge, CaWO<sub>4</sub>) Phase I (2015): 150 kg Full Scale: 500 - 1000 kg 10<sup>-10</sup> pb sensitivity

EDELWEISS, CRESST, Rosebud + others

LSM or possible extension





#### **Alliance for Astroparticle Physics**

New Alliance of Universities and Institutes in DE. Similar to MultiDark structure in Spain Affiliation with Institutes outside of DE (US: Kavli in Chi) funding possibilities exist within this framework 5 Year Funding Cycle - EURECA R&D